Exploring Black Holes, Second Edition

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Edwin F. Taylor, John Archibald Wheeler, and Edmund Bertschinger

For Instructors and Individual Readers

Exploring Black Holes: Introduction to General Relativity uses the properties of non-spinning and
 spinning black holes to introduce Albert Einstein's theory of curved spacetime and applies the resulting
 general relativity to the Universe around us. In this second edition -- labeled EBH2e -- coauthor
 Edmund Bertschinger joins Edwin F. Taylor to revise and expand the first edition by Edwin F. Taylor
 and John Archibald Wheeler published in 2000. What happened to the first edition? See the PS2 at the
 end of this document.

- 11 12 There is no p
- There is no published hard copy textbook of EBH2e. Instead, you may freely download the online
 version for personal and class use. Download one chapter or the entire book at the dropsite
 exploringblackholes.com
- 15
 16 THE STRATEGY: In EBH2e we choose to make every measurement and observation in a *local*17 inertial frame, so we can analyse them using special relativity. This assumes that almost everywhere -18 except possibly at the center of a black hole -- spacetime curves gently enough so that it is sufficiently
- 19 flat over a small region to use special relativity.
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- 21 Our strategy has two parts: The metric and the Principle of Maximal Aging. First the metric:
- Choose to index each event in curved spacetime with an arbitrarily chosen set of
 global coordinates, which we call *map coordinates*. Submit these map coordinates to
 Einstein's field equations, which return the *global metric*. The input to the global
 metric is the differential separation in global map coordinates between two adjacent
 events. The output of the global metric is the differential of a measureable quantity.
- One of several measurable quantities is the *wristwatch time*, the time between two
 events measured on a wristwatch that moves directly from one event to its neighbor.
- Next, we "elbow" the differentially small region of spacetime of the metric into a
 slightly larger region in which to construct a local inertial frame big enough to carry
 out our desired measurement. In this region the exact differential global metric
 becomes an approximate local metric that separates space and time for analysis by
 special relativity.
- Finally, the global metric allows us to connect measurements of, say, a fast-moving
 particle that passes along a *worldline* across one local inertial frame, then across a
 second local inertial frame far from the first.
- 41 42
- The second part of our strategy: The Principle of Maximal Aging, tells the stone how to move:
- 4344 The worldline of a free stone has maximum wristwatch time between adjacent
- 45 events. This leads to global constants of motion, such as map energy and map
- 46 angular momentum, which we use to predict global orbits.
- 47 This double strategy is diagrammed on the back cover of the book, reproduced on the following page.
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KEY IDEAS: Just three words summarize this book: <u>spacetime</u>, <u>motion</u>, <u>measurement</u>! The global metric--with arbitrary global coordinates--describes <u>spacetime</u>. The Principle of Maximal Aging describes free <u>motion</u>. Choose to report every <u>measurement</u> with respect to a local inertial frame.

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79 80 81	AAFrontMatter Constants and conversion factors. Fifty-seven questions.
82 83 84	AContents With a brief description of every chapter.
85 86 87	Chapter 1 Speeding Special relativity presented so as to lead naturally to general relativity
88 89 90	Chapter 2 The Bridge: SR to GR What is the difference between flat spacetime and curved spacetime?
91 92 93	Chapter 3 Curving Describe curved spacetime outside Earth and down to the center of a non-spinning black hole.
94 95 96	Chapter 4 The Global Positioning System (GPS) Locate yourself anywhere on Earth with a hand-held device which is useless without general relativity.
97 98	Chapter 5 Global and Local Metrics Over a sufficiently small region of spacetime, special relativity correctly describes every measurement.
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100	Chapter 6 Diving
101	Dive into a black hole. How do you feel? How long do you live? Why do you die?
102	Dive med a black noie. How do you leer. How long do you nve. Why do you die.
102	Chapter 7 Inside the Black Hole
103	A relayed life with spectacular effects around you and an ending certain
104	A relaxed file with spectacular criteris around you and an chung certain.
105	Chantar & Circular Orbita
100	What simpler "merking" orbits are evoluble to you meen a block hole?
10/	what circular parking orbits are available to you hear a black hole?
108	
109	Chapter 9 Orbiting
110	I ransfer between circular "parking" orbits. Close to the black hole, killer tides threaten humans and
	robots.
112	
113	Chapter 10 Advance of Mercury's Perihelion
114	An early victory of Einstein's brand new general relativity theory.
115	
116	Chapter 11 Orbits of Light
117	Trajectories of light around a black hole.
118	
119	Chapter 12 Diving Panoramas
120	What changing panoramas surround you as you drop into a black hole? What is the last thing you see?
121	
122	Chapter 13 Gravitational Mirages
123	Stars, galaxies, and black holes act like distorting lenses as we look outward into the Universe.
124	
125	Chapter 14 Expanding Universe
126	The shape in space and time of an expanding universe.
127	
128	Chapter 15 Cosmology
129	Ordinary matter? Dark matter? Dark energy? The Big Bang and its consequences.
130	
131	Chapter 16 Gravitational Waves
132	Gravitational waves tell us about cosmic catastrophes and regions close to the monster black hole.
133	
134	Chapter 17 Spinning Black Hole
135	Irresistible motion, multiple horizons, falling with a raindrop.
136	
137	Chapter 18 Circular Orbits around the Spinning Black Hole
138	Circular orbits far from and close to a spinning black hole. The accretion disk powers a quasar.
139	
140	Chapter 19 Orbiting the Spinning Black Hole
141	Moving between circular orbits near a spinning black hole
142	
143	Chapter 20 Orbits of Light around the Spinning Black Hole
144	What do you see as you orbit around and plunge into a spinning black hole
145	
146	Chapter 21 Travel Through the SpinningBH
147	Pass inward through the horizon of the spinning black hole; emerge into another Universe.
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149	Chapter 22 Deriving the Metric
150	(To be written) Einstein's field equations for static spherical spacetimes: derive the Schwarzschild
151	metric Metrics for spherical stars charged black holes white holes wormholes and inflationary
152	cosmology
153	cosmology.
155	V Annandix A Whaeler's Pules of Writing
154	Mativatal Simpliful Salf descriptive terminology, and the dullness of simply being
155	Mouvate! Simplify! Sen-descriptive terminology, and the dufiness of simply being.
150	
15/	W Appendix B Glossary
158	Words forbidden in our book. Definitions of permitted words.
159	
160	X Appendix C Acknowledgments
161	
162	Y Inside the Back Cover
163	General relativity briefing
164	
165	Z Back Cover
166	Diagrammed strategy of this book
167	
168	The spinless black hole is like a spinning black hole, but its gate to other universes is closed.
169	For the spinning black hole, this gate is ajar.
170	
171	Luc Longtin
172	
173	SOFTWARE GRothits
174	The interactive software program GRorbits by Slavomir Tuleia plots many figures in the book and is
175	available for the reader to plot orbits of a stone or a light flash around a black hole predicted by three
176	alternative theories.
177	
178	1 Newtonian mechanics
170	2 general relativity on a flat plane through the center of a non-spinning black hole
190	2. general relativity on the flat symmetry plane the equator of a spinning black hole with arbitrary
100	5. general relativity on the nat symmetry plane the equator of a <u>spinning</u> black hole, with arbitrary
101	mass and any value of spin permitted by the meory.
102	Here is the link to the CD white methods by http://stalais.org/anality/
183	Here is the link to the GRorbits website: <u>http://stuleja.org/grorbits/</u>
184	Here is the direct link to the GRorbits download site: <u>http://stuleja.org/grorbits/run.html</u>
185	
186	For Mac users, this puts GRorbits into the Applications folder.
187	
188	Please note, there are some pre-programmed scenarios which can be downloaded for use by the
189	GRorbits program, obtained from <u>http://stuleja.org/grorbits/scenarios.html</u>
190	
191	PS1. EXERCISES. Early chapters have lots of exercises. Later chapters have fewer. For exercise
192	assignments, the instructor may ask you to solve one or more QUERIES in the body of the chapter.
193	
194	PS2. WHAT HAPPENED TO THE FIRST EDITION of Exploring Black Holes by Edwin F. Taylor
195	and John Archibald Wheeler? The first edition is now out of print. The original publisher no longer
196	exists. Copies of the first edition are available online at exorbitant prices. In contrast, the contents of
197	this draft second edition, from this dropsite, contain more material, is more up to date, and is free.
198	

"Man fears time, but time fears the pyramids" by "The learned Arab of the twelfth century" (quot. 1923) who has not been identified. CORRECTIONS? OBJECTIONS? SUGGESTIONS? FOR EBH2e and this README file: Nandor Bokor at n bokor@yahoo.com born 1969 Edwin Taylor at eftaylor631@gmail.com born 1931 FOR the interactive software program GRorbits: Slavo Tuleja at stuleja@gmail.com John Archibald Wheeler 1911 - 2008 **ACKNOWLEDGMENTS** NANDOR BOKOR of the Department of Physics, Budapest University of Technology and Economics worked tirelessly on revisions of EBH2e and helped to prepare this online version. LUC LONGTIN, a professional legal French-English translator, made black holes an avocation -- a hobby. His detailed comments and suggestion on endless drafts of EBH2e corrected the physics and clarified the logic of presentation. JOHN ROGOSICH, President of Techsetters, Inc., contributed his services to design, redesign, and continually upgrade the LaTeX format of chapters under the demanding requirements of authors and reviewers. The LaTeX style file has the name EBH.sty Download file name: AAAAReadMe180422.pdf